Instrument Design Document
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## 1. Configuration Definition and Functional Description

Two NuSTAR optics modules focus X-rays over a broad energy range ( $6-80 \mathrm{keV}$ ) onto their respective CdZnTe detectors. Segmented, multilayer-coated mirrors are built up from a titanium base mandrel using graphite spacers and epoxy. The inner layers will have sextant segmentation (i.e. $\sim 60$ degree azimuth span mirrors) and the outer layer will have twelvetant segmentation (i.e. $\sim 30$ degree azimuth span mirrors). The transition from sextant to twelvetant mirrors will be made using three layers of substrates that will be multilayer coated and are expected to have good quality figure and thus may potentially provide extra effective area (over the baseline of 130 reflective layers). They will serve as an intermediate mandrel that will support the change in spacer configuration from the sextant to twelvetant geometry and will also provide overall structural support by tying together the sector sections. A skeletal outer mandrel will join the upper and lower halves of the optic. The optics modules will be co-aligned and mounted via Aluminum spider supports to canisters at the end of the extendable mast, which will be deployed after reaching orbit. The spider supports are designed to completely obscure the region between adjacent sextant segments. An illustration of a complete optics module is shown in Figure 1.

## 2. Coordinate System definition

The optical axis is well defined by our assembly approach in which each mirror segment is mounted to graphite spacers that are precisely machined to the correct angle and radius, which are measured with respect to fixed points on the inner optic mandrel. The optical axis will be confirmed with X-ray calibration measurements. Mechanical alignment points will be used to define a mechanical boresight that is co-aligned with the optical axis. While the exact rotational alignment about the optical axis is not critical to the optical imaging performance, it will be well defined based on the mechanical alignment points.


Figure 1: NuSTAR optics with spider mounts

## 3. Mechanical Description

The NuSTAR optics design is based on a conical approximation to a Wolter-I telescope. True Wolter-I telescopes are comprised of concentric shells of primary (parabolic) and secondary (hyperbolic) mirrors. They have been successfully used for the soft ( $\sim 0.1-10$ keV ) band starting with the Einstein Observatory to the Chandra and XMM-Newton Observatories of today. NuSTAR is designed for energies between 6 to 80 keV . To achieve high reflectivity at these energies, the nominal graze angles of the mirrors must be less than for softer X-rays; thus NuSTAR will have a relatively long focal length ( $\mathrm{F}=$ 10.150 m , defined as the length from the midpoint between the primary and secondary mirrors to the focal plane detector). Further, NuSTAR will employ depth graded multilayers to achieve enhanced, wide-band reflectivity (see Multilayer Instrument Design Document). The long focal length, together with the NuSTAR image performance requirement of 43 " half power diameter (HPD, the focal plane diameter within which half of the focused photons are enclosed) means that the primary and secondary mirrors can be approximated to be the surfaces of cones with minimal impact to the image performance. This significantly simplifies the optics design and manufacturing. The general telescope design parameters are given in Table 1.

### 3.1. Layer Geometry

NuStar will have two optics modules, each composed of 133 sets of conic shells, 3 of which will serve as the intermediate mandrel. The upper conic shell forms an angle $\alpha$ with respect to the optics axis, while the lower shell forms an angle $3 \alpha$ with respect to the
optics axis. In this way, on-axis photons are reflected from the lower shell surface by an angle $4 \alpha$ (for perfect mirrors).


Figure 2: Optics design. $F$ is the focal length, $L$ is the shell length, $\alpha$ is the graze angle, $g$ is the gap between the upper and lower shells and the radius is identified at several points.
Since the lengths of the shells, $L=225 \mathrm{~mm}$, is relatively small compared to the focal length, these conic shells are a good approximation to the parabolic and hyperbolic form for true Wolter I optics. Further, the angle $\alpha$ is small enough that the small angle approximation, $\sin (\alpha)=\tan (\alpha)=\alpha$, may be used for all calculations (at most it effects the calculations of the shell radii by just over 1 ppm ). In order to minimize the conic approximation aberration, the relationship between the angle $\alpha$ and the radius to the center of the upper shell, $r_{u o}$, is defined to be

$$
r_{u o}=4 \alpha F,
$$

where $F$ is the focal length as shown in Figure 2. To calculate the positions for all 130 sets of shells, it is convenient to use the top of the upper shell,

$$
r_{u t}=r_{u o}+\alpha L / 2 .
$$

The angle $\alpha$ can then be written as

$$
\begin{equation*}
\alpha=\left(r_{u t}\right) /(4 F+L / 2) \tag{1}
\end{equation*}
$$

The radii at other points of each layer are then given by

$$
\begin{aligned}
& r_{u b}=r_{u t}-\alpha L \\
& r_{l t}=r_{u t}-\alpha(L+4 g) \\
& r_{l b}=r_{u t}-4 \alpha(L+g)
\end{aligned}
$$

where $g=2 \mathrm{~mm}$ is the distance from the intersection of the upper and lower surfaces extended to both the upper bottom shell edge and the lower top shell edge (see Figure 2). A radius dependent gap between each shell is used such that the difference in radii $r_{u t}$ between consecutive shells $i$ and $i-1$ (where the shell $i-1$ has a smaller radius than the shell $i$ ) is

$$
\begin{equation*}
\Delta r_{u t}(i, i-1)=\alpha_{i} L f_{g a p}+t_{\text {glass }}, \tag{2}
\end{equation*}
$$

where $f_{\text {gap }}=1.3313$ is the variable gap parameter and $t_{\text {glass }}$ is the glass substrate thickness. Combining the last two equations and solving for a

$$
\alpha_{i}=\left(r_{u t, i-1}+\mathrm{t}_{\text {glass }}\right) /\left(4 \mathrm{~F}+\mathrm{L} / 2-\mathrm{Lf}_{\mathrm{gap}}\right)
$$

so that

$$
\begin{equation*}
r_{u t, i}=\left(r_{u t, i-1}+\mathrm{t}_{\text {glass }}\right)\left\{1+\left[(4 \mathrm{~F}) /\left(\mathrm{Lf}_{\text {gap }}\right)+1 /\left(2 \mathrm{f}_{\text {gap }}\right)-1\right]^{-1}\right\} . \tag{3}
\end{equation*}
$$

### 3.2. Mirror Segment Geometry

Each complete shell is comprised of individual segmented mirrors. The segmentation of the mirrors will change midway through the optic. The parameters for the inner and outer sets of shells are given in Table 2. The inner 65 shells will be built directly off the inner base mandrel with the upper and lower shells, each comprised of $\operatorname{six}\left(\mathrm{n}_{\text {segs }}=6\right), L=225$ mm long mirror segments made of $t_{\text {glass }}=0.21 \mathrm{~mm}$ thick, D263 glass substrates with multilayer coatings. Each mirror segment will be supported by five $\left(\mathrm{n}_{\text {spacers }}=5\right), \mathrm{w}_{\text {spacer }}=$ 1.2 mm wide graphite spacers that run the full axial length of the mirror segment (i.e., 225 mm long).
The transition from sextant to twelvetant mirrors will be made using three layers of substrates that will be multilayer coated and are expected to have good quality figure and thus may potentially provide extra effective area (over the baseline of 130 reflective layers). Based on an engineering analysis of stress at this interface, all three of these layers will be the standard thickness $(0.21 \mathrm{~mm})$ glass. These layers $(66-68)$ will serve as an intermediate mandrel that will support the change in spacer configuration from the sextant to twelvetant geometry and will also provide overall structural support by tying together the sector sections.

The next 65 shells are comprised of twelve ( $\mathrm{n}_{\text {segs }}=12$ ), 225 mm long mirror segments also made of 0.21 mm thick, D263 glass. Each mirror segment will also be supported by five $\left(\mathrm{n}_{\text {spacers }}=5\right), \mathrm{w}_{\text {spacer }}=1.2 \mathrm{~mm}$ wide graphite spacers that run the full axial length of the mirror segment (i.e., 225 mm long).
The mirror segments will be sized to allow for a $\mathrm{w}_{\text {gap }}=1.0 \mathrm{~mm}$ average gap between adjacent segments such that the arclength is given by

$$
\begin{equation*}
S=2 \pi r / \mathrm{n}_{\text {segs }}-\mathrm{w}_{\text {gap }}, \tag{4}
\end{equation*}
$$

where $r$ is the radius of the segment. The cut size of each mirror segment is thus defined by this equation with top (bottom) radius of each upper segment given by $r_{u t}\left(r_{u b}\right)$ and the top (bottom) radius of each lower segment given by $r_{l t}\left(r_{l b}\right)$ as defined by the equations in Section 3.1. These cut dimensions for the inner (outer) 65 layers are detailed in Table 8 (Table 9). The cut dimensions for the intermediate transition layers (66-69) are detailed in Table 7.

The 'skeletal' outer mandrel will join the upper and lower halves of the optic and tie the twelvetant sectors together azimuthally. The baseline engineering design is to tie the twelvetant sectors together at their adjacent outside spacers using strips of Titanium (8 mm wide and 1 mm thick) along the entire length of the optic ( 454 mm long from the top of the upper shell to the bottom of the lower shell).
Based on these design details, mass estimates are given in Table 3. The cut shell geometries are posted on the NuSTAR twiki as NuSTAR_CutShellGeometryV20.txt.

### 3.3. Spacer Configuration and Obscuration

The spacers are $\mathrm{w}_{\text {spacer }}=1.2 \mathrm{~mm}$ wide, except for up to the five innermost layers (1-5) where the spacers will be $\mathrm{w}_{\text {spacer }}=1.6 \mathrm{~mm}$ wide since this is the area of highest stress concentration. On average, a small epoxy fillet $\mathrm{w}_{\text {fillet }}=1.0 \mathrm{~mm}$ wide (i.e. 0.5 mm average on each side) around the spacer will cause added obscuration for each spacer. Thus, on average, each spacer is expected to obscure 2.2 mm of the segment. The aluminum spider supports that connect the inner optic mandrel to the canisters at the end of the spacecraft's extendable mast are designed to divide the optic into sextant sectors that align with the sextant segmented mirrors. In this way, the $\mathrm{n}_{\text {sector }}=6$ spider supports with $\mathrm{w}_{\text {spider }}=7.0 \mathrm{~mm}$ width will completely obscure the region between the outer spacers of adjacent sextant segments including the nominal epoxy fillet width. The spiders will similarly obscure this region on one side of each of the twelvetant sectors such that only a small portion ( 1.2 mm for each twelvetant) of the unsupported overhang region outside the outermost spacer will be unobscured the other side of each of the twelvetant sectors. Including the spider obscuration and gaps between shells (for the twelvetants portion not covered by the spiders), the total on-axis obscuration will be

$$
\begin{align*}
& \varepsilon_{\text {obscuration }}=\left[\mathrm{n}_{\text {segs }}\left(\mathrm{n}_{\text {spacers }}-2\right)\left(\mathrm{w}_{\text {spacer }}+\mathrm{w}_{\text {fillet }}\right)+\right. \\
& \left.\quad\left(\mathrm{n}_{\text {segs }}-\mathrm{n}_{\text {sectors }}\right)\left(\mathrm{w}_{\text {gap }}+2 \mathrm{~W}_{\text {spacer }}+2 \mathrm{~W}_{\text {fillet }}\right)+\mathrm{n}_{\text {sectors }} \mathrm{W}_{\text {spider }}\right] /\left(2 \pi r_{l b}\right), \tag{5}
\end{align*}
$$

where $n_{\text {spacers }}=5$ is the number of spacers, $n_{\text {segments }}$ is the number of mirror segments as previously defined (i.e., 6 for the inner and 12 for the outer layers), $\mathrm{n}_{\text {sector }}=6$ is the number of spider legs (i.e. primary sectors), $\mathrm{w}_{\text {spider }}=7.0 \mathrm{~mm}$ is the spider width and the epoxy fillet width ( $\mathrm{w}_{\text {fillet }}$ ) and spacer width ( $\mathrm{w}_{\text {spacer }}$ ) are as previously defined. The total obscuration for will range from $13-26 \%$. The central spacers are positioned such that they are positioned at $\varphi_{\mathrm{o}, i}=360^{\circ}(i+0.5) / \mathrm{n}_{\text {segs }}$ in azimuth coordinates where $i=1, \mathrm{n}_{\mathrm{segs}}$ is the $\mathrm{i}^{\text {th }}$ segment, and the gap between the $i$ and $i+1$ segments is at $\varphi_{\text {gap }, i}=360^{\circ}\left(i / \mathrm{n}_{\text {segs }}\right)$. The outside spacers will be positioned such that the center of the spacer is a tangential distance of $w_{\text {oh }}=2.4 \mathrm{~mm}$ from the gap position. The other two spacers will be positioned at $\Delta \varphi_{\mathrm{o}, i}=+-360^{\circ} / n_{\text {segs }} /\left(n_{\text {spacers }}-1\right)$ from the central spacer. This geometry is illustrated in Figure 3.

Table 1: Overall optics design parameters

| Parameter | Unit | Value | Uncertainty |
| :---: | :---: | :---: | :---: |
| Focal Length (F) | [mm] | $10150^{\text {a }}$ |  |
| Innermost Shell Radius ( $\mathrm{ruo}^{\text {) }}$ ) | [mm] | 54.418 |  |
| Outermost Shell Radius ( $\mathrm{r}_{\mathrm{u}}$ ) | [mm] | 191.221 |  |
| Shell Length (L) | [mm] | 225.0 |  |
| Half Gap btw Upper/Lower Shell (g) | [mm] | 2.0 |  |
| Epoxy Thickness (tepoxy) | [mm] | 0.002 |  |
| Variable Gap (fgap |  | $1.3313^{\text {a }}$ |  |
| PSF performance (HPD) | [arcsec] | 43 |  |
| Spacer Density ( $\rho_{\text {spacer }}$ ) | $\left[\mathrm{g} / \mathrm{cm}^{3}\right]$ | 1.88 | $\pm 0.09$ |
| Epoxy Density ( $\rho_{\text {epoxy }}$ ) | [ $\mathrm{g} / \mathrm{cm}^{3}$ ] | 1.16 |  |
| Epoxy Per Spacer Length | [mg/cm] | 2.5 | +0.5,-1.0 |
| Total Shells Per Module |  | 130+3 |  |
| Total Mounted Segments |  | 4680 |  |
| Total Cut/Coat Requirement |  | 5200 |  |
| Total Slumping Requirement |  | 4010 |  |
| Total Glass Sheet Requirement |  | 2999 |  |
| Total Spacers |  | 14256 |  |

${ }^{\text {a }}$ In early 2009, the focal length was changed from 10140 mm to 10150 mm and as a result, the Variable Gap parameter was changed from 1.33 to 1.3313 in order to maintain the same shell geometry (within 0.01 mm ).


Figure 3: Close-up of top of the optic.

Table 2: Optics design details

| Parameter | Unit | Inner | Interm. | Outer | Uncertainty |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Layers |  | 1-65 | 66-68 | 69-133 |  |
| Min Shell Radius ( $\mathrm{ruo}^{\text {) }}$ | [mm] | 54.418 | 105.479 | 108.480 | $\pm 0.003$ |
| Max Shell Radius ( $\mathrm{ruo}^{\text {) }}$ | [mm] | 104.494 | 107.472 | 191.221 | $\pm 0.003$ |
| Min Upper Shell Graze Angle ( $\alpha_{u}$ ) | [mrad] | 1.342 | 2.601 | 2.675 | $\pm 0.015$ |
| Max Upper Shell Graze Angle ( $\alpha_{u}$ ) | [mrad] | 2.576 | 2.650 | 4.715 | $\pm 0.015$ |
| Min Lower Shell Graze Angle ( $\alpha_{1}$ ) | [mrad] | 4.026 | 7.803 | 8.025 | $\pm 0.015$ |
| Max Lower Shell Graze Angle ( $\alpha_{1}$ ) | [mrad] | 7.728 | 7.950 | 14.145 | $\pm 0.015$ |
| Substrate Source |  | GSFC | GSFC | GSFC |  |
| Glass Type |  | D263 | D263 | D263 |  |
| Glass Density ( $\rho_{\text {glass }}$ ) | $\left[\mathrm{g} / \mathrm{cm}^{3}\right]$ | 2.51 | 2.51 | 2.51 | $\pm 0.01$ |
| Substrate Length ( $\mathrm{L}_{\text {glass }}$ ) | [mm] | 225.0 | 225.0 | 225.0 | $\pm 0.2$ |
| Shell Length (L) | [mm] | 225 | 225 | 225 |  |
| Sectors per (U/L) Shell ( $\mathrm{n}_{\text {seg }}$ ) |  | 6 | 6 | 12 |  |
| Gap btw Sectors ( $\mathrm{w}_{\text {gap }}$ ) | [mm] | 1.0 | 1.0 | 1.0 | +0.5,-1.0 |
| Sector Arclength Minimum | [mm] | 54.9 | 107.3 | 54.7 |  |
| Sector Arclength Maximum | [mm] | 107.7 | 111.9 | 99.4 |  |
| Sector Chord Length Minimum | [mm] | 52.6 | 102.7 | 54.1 |  |
| Sector Chord Length Maximum | [mm] | 103.9 | 106.9 | 98.3 |  |
| Total Mounted Segments |  | 1560 | 72 | 3120 |  |
| Spares per layer per telescope |  | 2 | 4 | 2 |  |
| Total Cut/Coat Requirement |  | 1820 | 96 | 3380 |  |
| Slump + Cut Yield |  | 90\% | 70\% | 85\% |  |
| Slump Pieces per Slump Mandrel |  | 1.0 | 1.0 | 2.0 |  |
| Total Slumping Requirement |  | 2022 | 137 | 1988 |  |
| Segments per Glass Sheet |  | 2.0 | 2.0 | 1.0 |  |
| Total Glass Sheet Requirement |  | 1011 | 69 | 1988 |  |
| Mirror Thickness ( $\mathrm{t}_{\text {glass }}$ ) | [mm] | 0.210 | 0.210 | 0.210 | $\pm 0.006$ |
| Spacers per Sector ( $\mathrm{n}_{\text {spacer }}$ ) |  | 5 | 10 | 5 |  |
| Spacer Width ( $\mathrm{w}_{\text {spacer }}$ ) | [mm] | $1.2^{\text {a }}$ | 1.2 | 1.2 | +0.0,-0.2 |
| Epoxy Fillet Width ( $\mathrm{w}_{\text {fillet }}$ ) | [mm] | 1.0 | 1.0 | 1.0 | +0.1,-0.3 |
| Spacer + Epoxy Width | [mm] | 2.2 | 2.6 | 2.2 |  |
| Distance btw Gap \& Spacer ( $\mathrm{w}_{\text {oh }}$ ) | [mm] | 2.4 | 2.4 | 2.4 |  |
| Max. Spacer/Gap Obscuration [\%] |  | 26\% | 22\% | 23\% |  |
| Min. Spacer/Gap Obscuration [\%] |  | 13\% | 23\% | 13\% |  |
| Total Spacers |  | 4752 | 360 | 9504 |  |
| Throughput Efficiency |  | 92\% ${ }^{\text {c }}$ | 0\% ${ }^{\text {b }}$ | 92\% ${ }^{\text {c }}$ | +5,-7\% |

${ }^{2} 1.6 \mathrm{~mm}$ for layers 1-5;
${ }^{\mathrm{b}}$ Intermediate layers will likely have significant reflectivity, but are not included in baseline calculations for effective area.
${ }^{\text {c }}$ The finite PSF of the actual mirrors will reduce the on-axis effective area by $3-7 \%$ and affect the off-axis effective area by $\pm 2 \%$. This however is dependent on the exact shape and energy dependence of the PSF and not solely the HPD, and thus a slightly more conservative throughput value is given here.

Table 3: Optics mass estimates per optic (note that the uncertainty for the total mass per optic is given here simply as the linear sum of all uncertainty terms).

| System | Unit | Estimate | Uncertainty |
| :--- | :--- | ---: | ---: |
| Glass Mirror Substrates | $[\mathrm{kg}]$ | 21.32 | $\pm 0.5$ |
| Graphite Spacers | $[\mathrm{kg}]$ | 5.48 | $+0.5,-1.0$ |
| Epoxy | $[\mathrm{kg}]$ | 0.67 | $+0.2,-0.3$ |
| Multilayer Coatings | $[\mathrm{kg}]$ | 0.56 | $+0.2,-0.0$ |
| Inner Base Mandrel | $[\mathrm{kg}$ | 2.85 | $+0.0,-0.3$ |
| Intermediate Mandrel | $[\mathrm{kg}$ | 0.68 | $+0.2,-0.3$ |
| Outer Mandrel | $[\mathrm{kg}$ | 0.40 | $+0.2,-0.1$ |
| ZSpider \& Mounts | $[\mathrm{kg}]$ | 5.00 | $+0.2,-0.4$ |
| Thermal | $[\mathrm{kg}]$ | 0.40 | $+0.3,-0.1$ |
| Total Mass per Optic | $[\mathrm{kg}]$ | 37.36 | $+2.3,-2.7$ |

### 3.4. Multilayer Design

### 3.4.1. Multilayer terminology

A multilayer is a stack of thin film of two different materials, deposited alternately on top of each other. To achieve the highest reflectivity the two materials must have a high density contrast and the materials are referred to as spacer, which is the low Z density material, and absorber, which is the high Z density material. One pair of spacer and absorber layers is called a bi-layer and has thickness $d$. The fraction of the absorber to the total thickness of the bi-layer is called $\Gamma=\mathrm{a}_{\mathrm{absorber}} / \mathrm{d}$. The bi-layer progression of thicknesses of the multilayer is given by the power-law $\mathrm{d}_{\mathrm{i}}=\mathrm{a} /(\mathrm{b}+\mathrm{i})^{\mathrm{c}}$, where $i$ runs from 1 to the total number of bi-layers, N , that are in the stack. The thinnest layers are closest to the substrate, and the thickest at the top, where 1 is the very top bi-layer.

The power-law depth grading described above is used for layers N to 2 . The gamma fraction, $\Gamma$, changes throughout the stack, starting at $\Gamma_{\text {high }}$ at the bottom of the stack and finishing at $\Gamma_{\text {low }}$ at the top.

The top layer, $\mathrm{N}=1$, is treated slightly differently, and is designed to most efficiently reflect energies below the critical energy, related to the critical angle of total external reflection. Because the critical angle of total external reflection is higher for the denser materials, the top layer is designed to resemble a monolithic slab of Pt or W , with only a thin SiC or Si layer to protect the Pt , and instead of using the $\Gamma$ of the stack (hereafter called $\Gamma_{\text {stack }}$ ), we use a different gamma, $\Gamma_{\text {top }}$, favoring a higher gamma. The bi-layer thickness used for the top layer is derived from the power-law formula, but the applied $\Gamma$ is different.

### 3.4.2. Design constrains and optical constants

The multilayer recipes have been optimized using the Figure of Merit optimization program (Peter Mao et al, 1999).
We have divided the 133 shells of the optics into 10 mirror groups. The optimized parameters are:
$\mathrm{D}_{\text {max }}$ - Maximum bi-layer thickness
c - power-law index
N - Total number of bi-layers

The materials selected for our multilayer are $\mathrm{Pt} / \mathrm{C}$ and $\mathrm{W} / \mathrm{Si}$. Due to coating restraints, $\Gamma$ cannot be less than 0.3 , and the total thickness of the stack cannot exceed 1.0 microns for $\mathrm{W} / \mathrm{Si}$ and 0.6 microns for $\mathrm{Pt} / \mathrm{C}$ and $\mathrm{Pt} / \mathrm{SiC}$.

The interfacial roughness for the materials are based on averages from experimental coatings, and are $\sigma_{\mathrm{P} t / \mathrm{C}}=4.5 \AA, \sigma_{\mathrm{W} / \mathrm{Si}}=4.3 \AA$

The optical constants have been calculated theoretically and been experimentally verified. The optical constant files used for these optimizations are:

$$
\begin{aligned}
& \text { SiC_NIST_CPJ_081013 } \\
& \text { C_NIST_CPJ_081013 } \\
& \text { Pt_NIST_CPJ_081013 } \\
& \text { Si_NIST_CPJ_081013 } \\
& \text { W_NIST_CPJ_081013 }
\end{aligned}
$$

The Figures below show the measured and theoretical curves for Pt. The theoretical curves were obtained from the National Institute of Standards and Technology.


### 3.4.3. The Flight Recipes

Table 4. shows the 10 optimized recipes for flight optics FM1 and FM2. Table 5. Shows the recipes for optic FM0. Group 1 represents the inner mirrors and group 10 the outer.

Table 4: Multilayer Recipes for FM1/FM2

| Group | Material | Angle [mrad] | Layer range | $\begin{aligned} & \mathrm{D}_{\text {min }} \\ & {[\AA \bar{\AA}]} \end{aligned}$ | $\begin{aligned} & \mathrm{D}_{\max } \\ & {[\AA]} \end{aligned}$ | N | c | $\begin{aligned} & \sigma \\ & {[\AA]} \end{aligned}$ | $\Gamma_{\text {Top }}$ | $\Gamma_{\text {high }}$ | $\begin{aligned} & \hline \mathrm{D}_{\text {stack }} \\ & {[\mu \mathrm{m}]} \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | Pt/C | $\begin{array}{r} 1.34- \\ 1.52 \end{array}$ | 1-12 | 29 | 133.66 | 145 | 0.245 | 4.5 | 0.7 | 0.45 | 0.558 |
| 2 | Pt/C | -1.73 | 13-24 | 29 | 131.58 | 145 | 0.228 | 4.5 | 0.7 | 0.45 | 0.547 |
| 3 | Pt/C | -1.96 | 25-36 | 29 | 129.56 | 145 | 0.234 | 4.5 | 0.7 | 0.45 | 0.551 |
| 4 | Pt/C | - 2.22 | 37-49 | 29 | 121.84 | 145 | 0.214 | 4.5 | 0.7 | 0.45 | 0.537 |
| 5 | Pt/C | -2.52 | 50-62 | 29 | 109.50 | 145 | 0.225 | 4.5 | 0.7 | 0.45 | 0.541 |
| 6 | Pt/C | -2.85 | 63-76 | 29 | 107.50 | 145 | 0.225 | 4.5 | 0.7 | 0.45 | 0.541 |
| 7 | Pt/C | -3.23 | 77-89 | 29 | 102.75 | 145 | 0.212 | 4.5 | 0.7 | 0.45 | 0.534 |
| 8 | W/Si | -3.67 | 90-104 | 25.0 | 95.228 | 291 | 0.238 | 4.3 | 0.8 | 0.38 | 0.955 |
| 9 | W/Si | -4.16 | 105-118 | 25.0 | 83.942 | 291 | 0.220 | 4.3 | 0.8 | 0.38 | 0.934 |
| 10 | W/Si | -4.72 | 119-133 | 25.0 | 74.471 | 291 | 0.190 | 4.3 | 0.8 | 0.38 | 0.902 |

Note that for groups $8-10, \Gamma_{\text {Top }}$ is equal to 0.8 . The $W$ thickness is thus $\Gamma_{\text {Toр }} * d_{\text {max }}$ but since $\left(1-\Gamma_{\text {Top }}\right) * d_{\text {max }}<25$ AA, the top Si layer has been fixed at 25 AA.

Table 5: Multilayer Recipes for FM0.

| Group | Materia I | Angle [mrad ] | Layer range | $\begin{aligned} & \mathrm{D}_{\text {min }} \\ & {[\AA \bar{\AA}]} \end{aligned}$ | $\begin{aligned} & \mathrm{D}_{\max } \\ & {[\AA]} \end{aligned}$ | N | c | $\begin{aligned} & \sigma \\ & {[\AA]} \end{aligned}$ | $\Gamma_{\text {Top }}$ | $\Gamma_{\text {high }}$ | $\begin{aligned} & D_{\text {stack }} \\ & {[\mu \mathrm{m}]} \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | Pt/SiC | $\begin{array}{r} 1.34- \\ 1.52 \end{array}$ | 1-12 | 31.7 | 128.125 | 150 | 0.245 | 4.5 | 0.7 | 0.45 | 0.631 |
| 2 | Pt/SiC | - 1.73 | 13-24 | 30.5 | 120.935 | 150 | 0.228 | 4.5 | 0.7 | 0.45 | 0.611 |
| 3 | Pt/SiC | - 1.96 | 25-36 | 29.0 | 119.500 | 155 | 0.234 | 4.5 | 0.7 | 0.45 | 0.592 |
| 4 | Pt/SiC | - 2.22 | 37-49 | 29.0 | 118.188 | 155 | 0.214 | 4.5 | 0.7 | 0.45 | 0.578 |
| 5 | Pt/SiC | - 2.52 | 50-62 | 29.0 | 107.750 | 155 | 0.225 | 4.5 | 0.7 | 0.45 | 0.575 |
| 6 | Pt/SiC | -2.85 | 63-76 | 29.0 | 103.250 | 155 | 0.225 | 4.5 | 0.7 | 0.45 | 0.586 |
| 7 | Pt/SiC | -3.23 | 77-89 | 29.0 | 98.750 | 155 | 0.212 | 4.5 | 0.7 | 0.45 | 0.577 |
| 8 | W/Si | -3.67 | 90-104 | 25.0 | 95.228 | 291 | 0.238 | 4.3 | 0.8 | 0.38 | 0.955 |
| 9 | W/Si | -4.16 | 105-118 | 25.0 | 83.942 | 291 | 0.220 | 4.3 | 0.8 | 0.38 | 0.934 |
| 10 | W/Si | -4.72 | 119-133 | 25.0 | 74.471 | 291 | 0.190 | 4.3 | 0.8 | 0.38 | 0.902 |

Note that for groups $8-10, \Gamma_{\text {Top }}$ is equal to 0.8 . The W thickness is thus $\Gamma_{\text {Top }} * d_{\text {max }}$ but since $\left(1-\Gamma_{\text {Top }}\right) * d_{\text {max }}<25$ AA, the top Si layer has been fixed at 25 AA.

## 4. Design Uncertainties

Based on the mechanical description in Section 3, the design uncertainties for relevant optics parameters are detailed in Table 1 and Table 2.

### 4.1. Spacers

Based on the vendor tolerance specification, the uncertainty in the spacer width is $\delta \mathrm{w}_{\text {spacer }}$ $= \pm 0.02 \mathrm{~mm}$ (i.e., $\sim 2 \%$ ). Six spacers from the first set of spacers we received from the vendor were measured to be within the tolerance specification with an average measured width of 1.22 mm and a measurement accuracy of $\pm 0.01 \mathrm{~mm}$.

Carbone-DS4 is the name of the graphite that is baselined for the spacers. It is manufactured by Carbone of America and machined to the required spacer size by Industrial Graphite. The graphite was chosen to closely match the CTE of the D263 glass substrates. The density has been measured to be $1.88 \pm 0.02 \mathrm{~g} / \mathrm{cm}^{3}$, based on 101 spacers measured in units of 10 to 11 spacers. Given that the density may vary from batch to batch we estimate the uncertainty in the graphite density to be $\pm 0.09 \mathrm{~g} / \mathrm{cm}^{3}$ (i.e. $5 \%$ ).

### 4.2. Epoxy

Since the epoxy is manually applied, we expect there will be variation in the individual fillet widths. The estimate of the average epoxy fillet ('squish-out') near the spacers, $\mathrm{w}_{\text {fillet }}=1.0 \mathrm{~mm}$ (i.e. 0.5 mm on each side of the spacer), is based of both X-ray and microscope measurements of the bondline width achieved with the baseline method of applying the EP-30 epoxy. We estimate that the uncertainty in the fillet width for NuSTAR will be $\delta \mathrm{w}_{\text {fillet }}=[+0.1 \mathrm{~mm},-0.3 \mathrm{~mm}]$.
Based on results from the epoxy development program, the current best estimate for the average mass of epoxy per unit glass/graphite bond length is $2.5 \mathrm{mg} / \mathrm{cm}$. The uncertainty is epoxy mass is estimated to be $+0.5,-1.0 \mathrm{mg} / \mathrm{cm}$.

### 4.3. Glass

The specification for the average gap between adjacent mirror segments ( $\mathrm{w}_{\mathrm{gap}}=1 \mathrm{~mm}$ gap) is based on the specification for the substrate cutting templates. Measurements of cut shells indicate that the average arclengths of the cut shells are slightly smaller than the cut specifications (i.e., leading to a slightly larger average gap), however, it is expected that this will only have a positive impact on the mass margin (i.e. less mass). Only one side of each twelvetant substrate will provide effective area near the gap between adjacent segments with the other side covered by the spider structure. Thus the twelvetants will be aligned to the 1 mm gap tolerance in the exposed region such that any extra gap in the region covered by the spider will not cause any decrease in effective area. The estimated uncertainty in the average gap of $+0.5 \mathrm{~mm},-1.0 \mathrm{~mm}$ provides sufficient tolerance in the substrate cut dimensions to assure high cut yields with no undesirable impact on the mass or effective area margins. Figure 4 and Figure 5 show the measured deviations from the cut geometry prescriptions for arc-length.

The uncertainty in average substrate length is expected to be $\delta L= \pm 0.2 \mathrm{~mm}$. The axial length was measured for a sample of 1448 substrates with $80 \%$ of them within the measurement accuracy of $\pm 0.25 \mathrm{~mm}$ and less than $7 \%$ with deviations of more than $\pm 0.75 \mathrm{~mm}$ from the design value of 225 mm .
The glass density $\left(2.51 \mathrm{~g} / \mathrm{cm}^{3}\right)$ and thickness $(0.210 \mathrm{~mm})$ is based on standard Schott D263 microsheets. While the density is expected to be very uniform with little uncertainty $\left(\sim 0.01 \mathrm{~g} / \mathrm{cm}^{3}\right)$, the uncertainty in the thickness is quoted by Schott to be $\pm 0.025 \mathrm{~mm}$. We have measured the thickness uniformity of individual sheets to be $\sim 0.003 \mathrm{~mm}$ throughout the sheet but slowly changing and $\sim<0.001 \mathrm{~mm}$ in the draw direction (which is chosen to be the axial slump direction). In terms of performance verification with surface metrology on the back surface of the mirrors, the thickness uncertainty poses a negligible contribution. Sheets from the same box ( 100 sheets per box) which come from the same glass production run show similarly low ( $<\sim 0.005 \mathrm{~mm}$ ) distribution in thickness. The thickness uncertainty quoted by Schott is a maximum deviation for any sheet from any production run. For the purpose of mass calculation, we are only concerned with the average thickness and density; thus, we estimate that the uncertainty in the mass estimate of the glass is $3 \%$. The thickness uncertainty will have an extremely small affect ( $\ll 1 \%$ ) on the off-axis effective area due to vignetting.
The mass of 23 flight mirror (from layers 1,9 and 27) was measured to be $-1.5 \pm 0.5$ relative to the specified mass based on the cut prescriptions and density and thickness specifications. While we do not have independent measurements of the glass thickness and density, these measurements indicate that density and thickness values specified in this document must be close to their actual values.

### 4.4. Structural

As previously mentioned, the spider supports will be aligned with to completely obscure the region between the outer spacers of adjacent sextant segments (and one side of each twelvetant segment) including the nominal epoxy fillet width. The tolerance of the spider width will be small $(< \pm 0.1 \mathrm{~mm})$. Thus the throughput near the spiders will only be negatively affected by excessive epoxy fillets ( $<1.0 \mathrm{~mm}$ ).

The baseline design for the structural hardware (e.g., inner, intermediate and outer mandrels; mounting spider and related components) is estimated to have a mass uncertainty of $+10 \%,-20 \%$.


Figure 4: Histogram of measured deviation of the mirror arc-length measurement from the cut shell specifications for 1448 mirrors. Top (Bot) refers to the arc-length measured at the top (bottom) of the mirror, 'Top-Bot' is the difference between the top and bottom measurement for each mirror, and min (max) is the minimum (maximum) value of the top and bottom measurements.


Figure 5: Integrated fraction of mirrors versus deviation from cut specification. See Figure 4 caption for description of data labels.

## 5. Current Best Estimate Performance

The effective area is calculated using a raytrace program that includes the optics geometry, structural obscuration (vignetting, spacers, epoxy fillet, etc.), multilayer reflectivity and a model of the expected scatter from the mirrors due to figure error. Based on the mechanical description in Section 3, the current best estimate of the effective area for each telescope coated with FM1/FM2 multilayer recipes, as a function
of energy and off-axis viewing angles 3 arcminutes and 6 arcminutes is plotted in Figure 6. The plot includes the reduction in effective area due to the thermal blankets and the aperture stop. The intermediate transition layers $66-68$ are expected to be comprised of multilayer coated mirrors that will provide additional effective area, but are not included in the effective area calculations.

The most recent response files can be found at
http://www.nustar.caltech.edu/for-astronomers/simulations


Figure 6: Effective Area as a function of energy for on- and off-axis photons for one telescope. Area includes reduction in area due to thermal blankets, the presence of aperture stop and scattering due to figure errors in the mirrors. The multilayer recipes used here are for FM1 and FM2.

Table 6: Mirror segment geometry for intermediate transition layers

| Layer | $\alpha_{u}$ | $\alpha_{1}$ | $r_{u t}$ | $r_{u b}$ | $r_{\text {lt }}$ | $r_{\text {lb }}$ | $n_{\text {segs }}$ | $W_{\text {gap }}$ | $S_{u t}$ | $S_{u b}$ | $s_{\text {lt }}$ | $s_{l b}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | [mrad] | [mrad] | [mm] | [mm] | [mm] | [mm] |  | [mm] | [mm] | [mm] | [mm] | [mm] |
| 66 | 2.598 | 7.794 | 105.772 | 105.188 | 105.167 | 103.413 | 6 | 1.0 | 109.76 | 109.15 | 109.13 | 107.29 |
| 67 | 2.622 | 7.866 | 106.768 | 106.178 | 106.157 | 104.387 | 6 | 1.0 | 110.81 | 110.19 | 110.17 | 108.31 |
| 68 | 2.647 | 7.941 | 107.771 | 107.175 | 107.154 | 105.367 | 6 | 1.0 | 111.86 | 111.23 | 111.21 | 109.34 |

Table 7: Mirror segment geometry for inner layers

| Layer | $\alpha_{u}$ | $\alpha_{1}$ | $r_{u t}$ | $r_{u b}$ | $r_{\text {lt }}$ | $r_{\text {lb }}$ | $n_{\text {segs }}$ | $W_{\text {gap }}$ | $S_{u t}$ | $S_{u b}$ | $S_{\text {lt }}$ | $s_{l b}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | [mrad] | [mrad] | [mm] | [mm] | [mm] | [mm] |  | [mm] | [mm] | [mm] | [mm] | [mm] |
| 1 | 1.340 | 4.020 | 54.569 | 54.267 | 54.257 | 53.352 | 6 | 1.0 | 56.14 | 55.83 | 55.82 | 54.87 |
| 2 | 1.355 | 4.065 | 55.185 | 54.881 | 54.870 | 53.955 | 6 | 1.0 | 56.79 | 56.47 | 56.46 | 55.50 |
| 3 | 1.371 | 4.113 | 55.805 | 55.497 | 55.486 | 54.560 | 6 | 1.0 | 57.44 | 57.12 | 57.10 | 56.14 |
| 4 | 1.386 | 4.158 | 56.431 | 56.119 | 56.108 | 55.172 | 6 | 1.0 | 58.09 | 57.77 | 57.76 | 56.78 |
| 5 | 1.402 | 4.206 | 57.061 | 56.745 | 56.734 | 55.788 | 6 | 1.0 | 58.75 | 58.42 | 58.41 | 57.42 |
| 6 | 1.417 | 4.251 | 57.695 | 57.377 | 57.365 | 56.409 | 6 | 1.0 | 59.42 | 59.08 | 59.07 | 58.07 |
| 7 | 1.433 | 4.299 | 58.334 | 58.012 | 58.000 | 57.033 | 6 | 1.0 | 60.09 | 59.75 | 59.74 | 58.72 |
| 8 | 1.449 | 4.347 | 58.978 | 58.652 | 58.640 | 57.662 | 6 | 1.0 | 60.76 | 60.42 | 60.41 | 59.38 |
| 9 | 1.465 | 4.395 | 59.627 | 59.297 | 59.285 | 58.297 | 6 | 1.0 | 61.44 | 61.10 | 61.08 | 60.05 |
| 10 | 1.481 | 4.443 | 60.281 | 59.947 | 59.936 | 58.936 | 6 | 1.0 | 62.13 | 61.78 | 61.76 | 60.72 |
| 11 | 1.497 | 4.491 | 60.938 | 60.602 | 60.590 | 59.579 | 6 | 1.0 | 62.81 | 62.46 | 62.45 | 61.39 |
| 12 | 1.513 | 4.539 | 61.602 | 61.262 | 61.250 | 60.228 | 6 | 1.0 | 63.51 | 63.15 | 63.14 | 62.07 |
| 13 | 1.530 | 4.590 | 62.270 | 61.926 | 61.914 | 60.881 | 6 | 1.0 | 64.21 | 63.85 | 63.84 | 62.75 |
| 14 | 1.546 | 4.638 | 62.943 | 62.595 | 62.583 | 61.539 | 6 | 1.0 | 64.91 | 64.55 | 64.54 | 63.44 |
| 15 | 1.563 | 4.689 | 63.622 | 63.270 | 63.258 | 62.203 | 6 | 1.0 | 65.62 | 65.26 | 65.24 | 64.14 |
| 16 | 1.579 | 4.737 | 64.305 | 63.949 | 63.937 | 62.871 | 6 | 1.0 | 66.34 | 65.97 | 65.95 | 64.84 |
| 17 | 1.596 | 4.788 | 64.993 | 64.633 | 64.621 | 63.543 | 6 | 1.0 | 67.06 | 66.68 | 66.67 | 65.54 |
| 18 | 1.613 | 4.839 | 65.685 | 65.323 | 65.310 | 64.221 | 6 | 1.0 | 67.79 | 67.41 | 67.39 | 66.25 |
| 19 | 1.631 | 4.893 | 66.384 | 66.018 | 66.004 | 64.904 | 6 | 1.0 | 68.52 | 68.13 | 68.12 | 66.97 |
| 20 | 1.648 | 4.944 | 67.088 | 66.718 | 66.704 | 65.592 | 6 | 1.0 | 69.25 | 68.87 | 68.85 | 67.69 |
| 21 | 1.665 | 4.995 | 67.796 | 67.422 | 67.408 | 66.284 | 6 | 1.0 | 70.00 | 69.60 | 69.59 | 68.41 |
| 22 | 1.683 | 5.049 | 68.511 | 68.133 | 68.119 | 66.983 | 6 | 1.0 | 70.74 | 70.35 | 70.33 | 69.14 |
| 23 | 1.700 | 5.100 | 69.230 | 68.848 | 68.834 | 67.687 | 6 | 1.0 | 71.50 | 71.10 | 71.08 | 69.88 |
| 24 | 1.718 | 5.154 | 69.955 | 69.569 | 69.555 | 68.395 | 6 | 1.0 | 72.26 | 71.85 | 71.84 | 70.62 |
| 25 | 1.736 | 5.208 | 70.685 | 70.295 | 70.281 | 69.109 | 6 | 1.0 | 73.02 | 72.61 | 72.60 | 71.37 |
| 26 | 1.754 | 5.262 | 71.420 | 71.026 | 71.012 | 69.828 | 6 | 1.0 | 73.79 | 73.38 | 73.36 | 72.12 |
| 27 | 1.772 | 5.316 | 72.161 | 71.763 | 71.748 | 70.552 | 6 | 1.0 | 74.57 | 74.15 | 74.13 | 72.88 |
| 28 | 1.791 | 5.373 | 72.907 | 72.505 | 72.490 | 71.281 | 6 | 1.0 | 75.35 | 74.93 | 74.91 | 73.65 |
| 29 | 1.809 | 5.427 | 73.660 | 73.252 | 73.238 | 72.017 | 6 | 1.0 | 76.14 | 75.71 | 75.69 | 74.42 |
| 30 | 1.828 | 5.484 | 74.418 | 74.006 | 73.992 | 72.758 | 6 | 1.0 | 76.93 | 76.50 | 76.48 | 75.19 |
| 31 | 1.847 | 5.541 | 75.181 | 74.765 | 74.750 | 73.504 | 6 | 1.0 | 77.73 | 77.29 | 77.28 | 75.97 |
| 32 | 1.866 | 5.598 | 75.949 | 75.529 | 75.514 | 74.255 | 6 | 1.0 | 78.53 | 78.09 | 78.08 | 76.76 |
| 33 | 1.885 | 5.655 | 76.724 | 76.300 | 76.285 | 75.012 | 6 | 1.0 | 79.35 | 78.90 | 78.89 | 77.55 |
| 34 | 1.904 | 5.712 | 77.504 | 77.076 | 77.061 | 75.775 | 6 | 1.0 | 80.16 | 79.71 | 79.70 | 78.35 |
| 35 | 1.923 | 5.769 | 78.290 | 77.858 | 77.842 | 76.544 | 6 | 1.0 | 80.99 | 80.53 | 80.52 | 79.16 |
| 36 | 1.942 | 5.826 | 79.081 | 78.645 | 78.629 | 77.318 | 6 | 1.0 | 81.81 | 81.36 | 81.34 | 79.97 |
| 37 | 1.962 | 5.886 | 79.880 | 79.438 | 79.423 | 78.098 | 6 | 1.0 | 82.65 | 82.19 | 82.17 | 80.78 |
| 38 | 1.982 | 5.946 | 80.683 | 80.237 | 80.221 | 78.883 | 6 | 1.0 | 83.49 | 83.02 | 83.01 | 81.61 |
| 39 | 2.002 | 6.006 | 81.492 | 81.042 | 81.026 | 79.674 | 6 | 1.0 | 84.34 | 83.87 | 83.85 | 82.43 |
| 40 | 2.022 | 6.066 | 82.308 | 81.854 | 81.837 | 80.472 | 6 | 1.0 | 85.19 | 84.72 | 84.70 | 83.27 |
| 41 | 2.042 | 6.126 | 83.130 | 82.670 | 82.654 | 81.276 | 6 | 1.0 | 86.05 | 85.57 | 85.56 | 84.11 |
| 42 | 2.062 | 6.186 | 83.958 | 83.494 | 83.478 | 82.086 | 6 | 1.0 | 86.92 | 86.43 | 86.42 | 84.96 |
| 43 | 2.083 | 6.249 | 84.791 | 84.323 | 84.306 | 82.900 | 6 | 1.0 | 87.79 | 87.30 | 87.29 | 85.81 |
| 44 | 2.103 | 6.309 | 85.632 | 85.158 | 85.142 | 83.722 | 6 | 1.0 | 88.67 | 88.18 | 88.16 | 86.67 |
| 45 | 2.124 | 6.372 | 86.478 | 86.000 | 85.983 | 84.549 | 6 | 1.0 | 89.56 | 89.06 | 89.04 | 87.54 |
| 46 | 2.145 | 6.435 | 87.330 | 86.848 | 86.831 | 85.383 | 6 | 1.0 | 90.45 | 89.95 | 89.93 | 88.41 |
| 47 | 2.166 | 6.498 | 88.189 | 87.701 | 87.684 | 86.222 | 6 | 1.0 | 91.35 | 90.84 | 90.82 | 89.29 |
| 48 | 2.187 | 6.561 | 89.054 | 88.562 | 88.544 | 87.068 | 6 | 1.0 | 92.26 | 91.74 | 91.72 | 90.18 |
| 49 | 2.209 | 6.627 | 89.926 | 89.428 | 89.411 | 87.920 | 6 | 1.0 | 93.17 | 92.65 | 92.63 | 91.07 |
| 50 | 2.230 | 6.690 | 90.804 | 90.302 | 90.284 | 88.779 | 6 | 1.0 | 94.09 | 93.56 | 93.55 | 91.97 |
| 51 | 2.252 | 6.756 | 91.688 | 91.182 | 91.164 | 89.644 | 6 | 1.0 | 95.02 | 94.49 | 94.47 | 92.87 |
| 52 | 2.274 | 6.822 | 92.580 | 92.068 | 92.050 | 90.515 | 6 | 1.0 | 95.95 | 95.41 | 95.39 | 93.79 |
| 53 | 2.296 | 6.888 | 93.477 | 92.961 | 92.942 | 91.393 | 6 | 1.0 | 96.89 | 96.35 | 96.33 | 94.71 |
| 54 | 2.318 | 6.954 | 94.382 | 93.860 | 93.842 | 92.277 | 6 | 1.0 | 97.84 | 97.29 | 97.27 | 95.63 |
| 55 | 2.341 | 7.023 | 95.293 | 94.767 | 94.748 | 93.168 | 6 | 1.0 | 98.79 | 98.24 | 98.22 | 96.57 |
| 56 | 2.363 | 7.089 | 96.211 | 95.679 | 95.660 | 94.065 | 6 | 1.0 | 99.75 | 99.19 | 99.18 | 97.50 |
| 57 | 2.386 | 7.158 | 97.135 | 96.599 | 96.579 | 94.969 | 6 | 1.0 | 100.72 | 100.16 | 100.14 | 98.45 |
| 58 | 2.409 | 7.227 | 98.067 | 97.525 | 97.506 | 95.880 | 6 | 1.0 | 101.70 | 101.13 | 101.11 | 99.40 |
| 59 | 2.432 | 7.296 | 99.006 | 98.458 | 98.439 | 96.797 | 6 | 1.0 | 102.68 | 102.11 | 102.09 | 100.37 |
| 60 | 2.455 | 7.365 | 99.951 | 99.399 | 99.379 | 97.722 | 6 | 1.0 | 103.67 | 103.09 | 103.07 | 101.33 |
| 61 | 2.478 | 7.434 | 100.904 | 100.346 | 100.326 | 98.654 | 6 | 1.0 | 104.67 | 104.08 | 104.06 | 102.31 |
| 62 | 2.502 | 7.506 | 101.862 | 101.300 | 101.280 | 99.591 | 6 | 1.0 | 105.67 | 105.08 | 105.06 | 103.29 |
| 63 | 2.526 | 7.578 | 102.829 | 102.261 | 102.241 | 100.536 | 6 | 1.0 | 106.68 | 106.09 | 106.07 | 104.28 |
| 64 | 2.550 | 7.650 | 103.803 | 103.229 | 103.209 | 101.487 | 6 | 1.0 | 107.70 | 107.10 | 107.08 | 105.28 |
| 65 | 2.574 | 7.722 | 104.784 | 104.204 | 104.184 | 102.446 | 6 | 1.0 | 108.73 | 108.12 | 108.10 | 106.28 |

Table 8: Mirror segment geometry for outer layers

| Layer | $\alpha_{u}$ | $\alpha_{1}$ | $r_{u t}$ | $r_{u b}$ | $r_{\text {lt }}$ | $r_{l b}$ | $n_{\text {segs }}$ | $W_{\text {gap }}$ | $S_{u t}$ | $S_{u b}$ | $S_{\text {lt }}$ | $S_{l b}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | [mrad] | [mrad] | [mm] | [mm] | [mm] | [mm] |  | [mm] | [mm] | [mm] | [mm] | [mm] |
| 69 | 2.672 | 8.016 | 108.781 | 108.179 | 108.158 | 106.354 | 12 | 1.0 | 55.96 | 55.64 | 55.63 | 54.69 |
| 70 | 2.697 | 8.091 | 109.799 | 109.193 | 109.171 | 107.351 | 12 | 1.0 | 56.49 | 56.17 | 56.16 | 55.21 |
| 71 | 2.722 | 8.166 | 110.824 | 110.212 | 110.190 | 108.353 | 12 | 1.0 | 57.03 | 56.71 | 56.70 | 55.73 |
| 72 | 2.747 | 8.241 | 111.857 | 111.239 | 111.217 | 109.363 | 12 | 1.0 | 57.57 | 57.24 | 57.23 | 56.26 |
| 73 | 2.773 | 8.319 | 112.898 | 112.274 | 112.252 | 110.380 | 12 | 1.0 | 58.11 | 57.79 | 57.77 | 56.79 |
| 74 | 2.799 | 8.397 | 113.946 | 113.316 | 113.294 | 111.404 | 12 | 1.0 | 58.66 | 58.33 | 58.32 | 57.33 |
| 75 | 2.825 | 8.475 | 115.003 | 114.367 | 114.345 | 112.438 | 12 | 1.0 | 59.22 | 58.88 | 58.87 | 57.87 |
| 76 | 2.851 | 8.553 | 116.067 | 115.425 | 115.402 | 113.478 | 12 | 1.0 | 59.77 | 59.44 | 59.42 | 58.42 |
| 77 | 2.877 | 8.631 | 117.139 | 116.491 | 116.468 | 114.526 | 12 | 1.0 | 60.33 | 59.99 | 59.98 | 58.97 |
| 78 | 2.904 | 8.712 | 118.218 | 117.564 | 117.541 | 115.581 | 12 | 1.0 | 60.90 | 60.56 | 60.54 | 59.52 |
| 79 | 2.930 | 8.790 | 119.306 | 118.646 | 118.623 | 116.645 | 12 | 1.0 | 61.47 | 61.12 | 61.11 | 60.08 |
| 80 | 2.957 | 8.871 | 120.402 | 119.736 | 119.713 | 117.717 | 12 | 1.0 | 62.04 | 61.69 | 61.68 | 60.64 |
| 81 | 2.984 | 8.952 | 121.506 | 120.834 | 120.810 | 118.796 | 12 | 1.0 | 62.62 | 62.27 | 62.26 | 61.20 |
| 82 | 3.012 | 9.036 | 122.618 | 121.940 | 121.916 | 119.883 | 12 | 1.0 | 63.20 | 62.85 | 62.84 | 61.77 |
| 83 | 3.039 | 9.117 | 123.738 | 123.054 | 123.030 | 120.978 | 12 | 1.0 | 63.79 | 63.43 | 63.42 | 62.34 |
| 84 | 3.067 | 9.201 | 124.867 | 124.177 | 124.152 | 122.082 | 12 | 1.0 | 64.38 | 64.02 | 64.01 | 62.92 |
| 85 | 3.095 | 9.285 | 126.004 | 125.308 | 125.283 | 123.194 | 12 | 1.0 | 64.98 | 64.61 | 64.60 | 63.50 |
| 86 | 3.123 | 9.369 | 127.149 | 126.447 | 126.422 | 124.314 | 12 | 1.0 | 65.58 | 65.21 | 65.19 | 64.09 |
| 87 | 3.151 | 9.453 | 128.303 | 127.595 | 127.569 | 125.442 | 12 | 1.0 | 66.18 | 65.81 | 65.80 | 64.68 |
| 88 | 3.180 | 9.540 | 129.466 | 128.750 | 128.725 | 126.578 | 12 | 1.0 | 66.79 | 66.41 | 66.40 | 65.28 |
| 89 | 3.209 | 9.627 | 130.637 | 129.915 | 129.889 | 127.723 | 12 | 1.0 | 67.40 | 67.02 | 67.01 | 65.88 |
| 90 | 3.238 | 9.714 | 131.817 | 131.089 | 131.063 | 128.877 | 12 | 1.0 | 68.02 | 67.64 | 67.62 | 66.48 |
| 91 | 3.267 | 9.801 | 133.006 | 132.270 | 132.244 | 130.039 | 12 | 1.0 | 68.64 | 68.26 | 68.24 | 67.09 |
| 92 | 3.296 | 9.888 | 134.203 | 133.461 | 133.435 | 131.210 | 12 | 1.0 | 69.27 | 68.88 | 68.87 | 67.70 |
| 93 | 3.326 | 9.978 | 135.409 | 134.661 | 134.634 | 132.389 | 12 | 1.0 | 69.90 | 69.51 | 69.49 | 68.32 |
| 94 | 3.356 | 10.068 | 136.625 | 135.869 | 135.843 | 133.577 | 12 | 1.0 | 70.54 | 70.14 | 70.13 | 68.94 |
| 95 | 3.386 | 10.158 | 137.849 | 137.087 | 137.060 | 134.774 | 12 | 1.0 | 71.18 | 70.78 | 70.76 | 69.57 |
| 96 | 3.416 | 10.248 | 139.082 | 138.314 | 138.286 | 135.981 | 12 | 1.0 | 71.82 | 71.42 | 71.41 | 70.20 |
| 97 | 3.447 | 10.341 | 140.325 | 139.549 | 139.522 | 137.195 | 12 | 1.0 | 72.47 | 72.07 | 72.05 | 70.84 |
| 98 | 3.477 | 10.431 | 141.576 | 140.794 | 140.766 | 138.419 | 12 | 1.0 | 73.13 | 72.72 | 72.70 | 71.48 |
| 99 | 3.508 | 10.524 | 142.837 | 142.047 | 142.019 | 139.651 | 12 | 1.0 | 73.79 | 73.38 | 73.36 | 72.12 |
| 100 | 3.540 | 10.620 | 144.107 | 143.311 | 143.282 | 140.893 | 12 | 1.0 | 74.45 | 74.04 | 74.02 | 72.77 |
| 101 | 3.571 | 10.713 | 145.387 | 144.583 | 144.555 | 142.144 | 12 | 1.0 | 75.12 | 74.70 | 74.69 | 73.43 |
| 102 | 3.603 | 10.809 | 146.676 | 145.866 | 145.837 | 143.405 | 12 | 1.0 | 75.80 | 75.38 | 75.36 | 74.09 |
| 103 | 3.635 | 10.905 | 147.975 | 147.157 | 147.128 | 144.674 | 12 | 1.0 | 76.48 | 76.05 | 76.04 | 74.75 |
| 104 | 3.667 | 11.001 | 149.284 | 148.458 | 148.429 | 145.954 | 12 | 1.0 | 77.16 | 76.73 | 76.72 | 75.42 |
| 105 | 3.699 | 11.097 | 150.601 | 149.769 | 149.739 | 147.242 | 12 | 1.0 | 77.85 | 77.42 | 77.40 | 76.10 |
| 106 | 3.732 | 11.196 | 151.929 | 151.089 | 151.059 | 148.540 | 12 | 1.0 | 78.55 | 78.11 | 78.09 | 76.78 |
| 107 | 3.765 | 11.295 | 153.267 | 152.419 | 152.389 | 149.848 | 12 | 1.0 | 79.25 | 78.81 | 78.79 | 77.46 |
| 108 | 3.798 | 11.394 | 154.614 | 153.760 | 153.729 | 151.166 | 12 | 1.0 | 79.96 | 79.51 | 79.49 | 78.15 |
| 109 | 3.831 | 11.493 | 155.972 | 155.110 | 155.079 | 152.493 | 12 | 1.0 | 80.67 | 80.22 | 80.20 | 78.85 |
| 110 | 3.865 | 11.595 | 157.340 | 156.470 | 156.439 | 153.830 | 12 | 1.0 | 81.38 | 80.93 | 80.91 | 79.55 |
| 111 | 3.898 | 11.694 | 158.718 | 157.840 | 157.809 | 155.178 | 12 | 1.0 | 82.10 | 81.65 | 81.63 | 80.25 |
| 112 | 3.933 | 11.799 | 160.105 | 159.221 | 159.189 | 156.534 | 12 | 1.0 | 82.83 | 82.37 | 82.35 | 80.96 |
| 113 | 3.967 | 11.901 | 161.503 | 160.611 | 160.579 | 157.901 | 12 | 1.0 | 83.56 | 83.10 | 83.08 | 81.68 |
| 114 | 4.002 | 12.006 | 162.912 | 162.012 | 161.980 | 159.278 | 12 | 1.0 | 84.30 | 83.83 | 83.81 | 82.40 |
| 115 | 4.036 | 12.108 | 164.331 | 163.423 | 163.391 | 160.666 | 12 | 1.0 | 85.04 | 84.57 | 84.55 | 83.12 |
| 116 | 4.071 | 12.213 | 165.761 | 164.845 | 164.812 | 162.065 | 12 | 1.0 | 85.79 | 85.31 | 85.30 | 83.86 |
| 117 | 4.107 | 12.321 | 167.201 | 166.277 | 166.244 | 163.472 | 12 | 1.0 | 86.55 | 86.06 | 86.05 | 84.59 |
| 118 | 4.143 | 12.429 | 168.652 | 167.720 | 167.687 | 164.890 | 12 | 1.0 | 87.31 | 86.82 | 86.80 | 85.34 |
| 119 | 4.178 | 12.534 | 170.114 | 169.174 | 169.141 | 166.320 | 12 | 1.0 | 88.07 | 87.58 | 87.56 | 86.09 |
| 120 | 4.215 | 12.645 | 171.586 | 170.638 | 170.604 | 167.759 | 12 | 1.0 | 88.84 | 88.35 | 88.33 | 86.84 |
| 121 | 4.251 | 12.753 | 173.069 | 172.113 | 172.079 | 169.209 | 12 | 1.0 | 89.62 | 89.12 | 89.10 | 87.60 |
| 122 | 4.288 | 12.864 | 174.563 | 173.599 | 173.564 | 170.670 | 12 | 1.0 | 90.40 | 89.90 | 89.88 | 88.36 |
| 123 | 4.325 | 12.975 | 176.070 | 175.096 | 175.062 | 172.142 | 12 | 1.0 | 91.19 | 90.68 | 90.66 | 89.13 |
| 124 | 4.362 | 13.086 | 177.586 | 176.604 | 176.569 | 173.625 | 12 | 1.0 | 91.98 | 91.47 | 91.45 | 89.91 |
| 125 | 4.399 | 13.197 | 179.114 | 178.124 | 178.089 | 175.120 | 12 | 1.0 | 92.78 | 92.27 | 92.25 | 90.69 |
| 126 | 4.437 | 13.311 | 180.653 | 179.655 | 179.619 | 176.624 | 12 | 1.0 | 93.59 | 93.07 | 93.05 | 91.48 |
| 127 | 4.475 | 13.425 | 182.203 | 181.197 | 181.161 | 178.140 | 12 | 1.0 | 94.40 | 93.87 | 93.86 | 92.27 |
| 128 | 4.514 | 13.542 | 183.766 | 182.750 | 182.714 | 179.667 | 12 | 1.0 | 95.22 | 94.69 | 94.67 | 93.07 |
| 129 | 4.552 | 13.656 | 185.339 | 184.315 | 184.278 | 181.206 | 12 | 1.0 | 96.04 | 95.51 | 95.49 | 93.88 |
| 130 | 4.591 | 13.773 | 186.924 | 185.892 | 185.855 | 182.756 | 12 | 1.0 | 96.87 | 96.33 | 96.31 | 94.69 |
| 131 | 4.631 | 13.893 | 188.521 | 187.479 | 187.442 | 184.316 | 12 | 1.0 | 97.71 | 97.16 | 97.14 | 95.51 |
| 132 | 4.670 | 14.010 | 190.130 | 189.080 | 189.042 | 185.890 | 12 | 1.0 | 98.55 | 98.00 | 97.98 | 96.33 |
| 133 | 4.710 | 14.130 | 191.751 | 190.691 | 190.653 | 187.474 | 12 | 1.0 | 99.40 | 98.85 | 98.83 | 97.16 |

